

ON THE STABILITY OF THE FLOW OF A VISCOELASTIC FLUID DOWN AN INCLINED PLANE

A. T. Listrov

Zhurnal Prikladnoi Mekhaniki i Tekhnicheskoi Fiziki, No. 5, pp. 102-105, 1965

ABSTRACT: The paper presents an analysis of laminar flow of a film of viscoelastic fluid flowing under gravity down an infinite inclined plane. It is assumed that the mechanical behavior of the fluid can be represented by a generalized Maxwell model, whose constitutive equation contains a time derivative of the deviator of the stress tensor in the Jaumann sense [1, 2]. The equations of motion of the viscoelastic fluid considered here admit an exact solution for the case of rectilinear laminar flow with a plane free boundary. The stability of this flow with respect to surface waves is investigated by the method of successive approximations described in [3, 4].

1. The viscoelastic flow of several real fluids can be approximately described by the generalized Maxwell model [2]

$$\begin{aligned} \epsilon_{ii} = 0, \quad s_{ik} = 2\mu\epsilon_{ik} - \lambda \left(\frac{\partial s_{ik}}{\partial t^*} + v_j \frac{\partial s_{ik}}{\partial x_j} + \omega_{ij}s_{jk} + \omega_{kj}s_{ij} \right) \quad (1.1) \\ \epsilon_{ik} = \frac{1}{2} \left(\frac{\partial v_k}{\partial x_i} + \frac{\partial v_i}{\partial x_k} \right), \quad \omega_{ik} = \frac{1}{2} \left(\frac{\partial v_k}{\partial x_i} - \frac{\partial v_i}{\partial x_k} \right) \quad (i, j, k = 1, 2, 3). \end{aligned}$$

Here s_{jk} are the components of the deviator of the stress tensor, v_j are the components of the velocity vector, μ is the coefficient of viscosity, λ is the relaxation time, and x_j are rectangular Cartesian coordinates.

The viscoelastic flow of a relaxing fluid with finite rates of strain is governed by (1.1) and the equations of motion

$$\rho \left(\frac{\partial v_k}{\partial t^*} + \frac{\partial v_k}{\partial x_j} v_j \right) = - \frac{\partial p^*}{\partial x_k} + \frac{\partial s_{kj}}{\partial x_j} + F_k \quad (j, k = 1, 2, 3). \quad (1.2)$$

Here F_k are the projections of the body force, ρ is the density, and p^* is the hydrostatic pressure. To analyze the stability of the flow of a film of thickness d flowing under the action of gravity down a plane inclined at an angle β with respect to the horizontal, we introduce the dimensionless variables

$$\begin{aligned} t = \frac{t^*k}{d}, \quad p = \frac{p^*}{\rho k^2}, \quad x = \frac{x_1}{d}, \quad u = \frac{v_1}{k}, \quad v = \frac{v_2}{k} \\ y = \frac{x_2}{d}, \quad s_{xx} = \frac{s_{11}}{\rho k^2}, \quad s_{yy} = \frac{s_{22}}{\rho k^2}, \quad s_{xy} = \frac{s_{12}}{\rho k^2}, \quad k = \frac{\rho g d^2 \sin \beta}{3\mu} \end{aligned}$$

where g is the acceleration of gravity.

In the case of two-dimensional unsteady flow with $F_1 = \rho g \sin \beta$, $F_2 = \rho g \cos \beta$, $F_3 = 0$, Eqs. (1.1), (1.2) take the form

$$\begin{aligned} \frac{\partial u}{\partial t} + \frac{\partial u}{\partial x} u + \frac{\partial u}{\partial y} v = - \frac{\partial p}{\partial x} + \frac{3}{R} + \frac{\partial s_{xx}}{\partial x} + \frac{\partial s_{xy}}{\partial y}, \\ \frac{\partial v}{\partial t} + \frac{\partial v}{\partial x} u + \frac{\partial v}{\partial y} v = - \frac{\partial p}{\partial y} + \frac{3 \operatorname{ctg} \beta}{R} + \frac{\partial s_{xy}}{\partial x} + \frac{\partial s_{yy}}{\partial y}, \\ s_{xx} = -s_{yy} = \frac{2}{R} \frac{\partial u}{\partial x} - \tau \left[\frac{\partial s_{xx}}{\partial t} + \frac{\partial s_{xx}}{\partial x} u + \frac{\partial s_{xx}}{\partial y} v + s_{xy} \left(\frac{\partial u}{\partial y} - \frac{\partial v}{\partial x} \right) \right], \\ s_{xy} = \frac{1}{R} \left(\frac{\partial u}{\partial y} + \frac{\partial v}{\partial x} \right) - \tau \left[\frac{\partial s_{xy}}{\partial t} + \frac{\partial s_{xy}}{\partial x} u + \frac{\partial s_{xy}}{\partial y} v + s_{xx} \left(\frac{\partial u}{\partial x} - \frac{\partial v}{\partial y} \right) \right], \\ \frac{\partial u}{\partial x} + \frac{\partial v}{\partial y} = 0 \quad \left(R = \frac{k d \rho}{\mu}, \quad \tau = \frac{\lambda k}{d} \right). \quad (1.3) \end{aligned}$$

These equations are satisfied by the relations

$$\begin{aligned} s_{xy}^0 = - \frac{3}{R} y, \quad s_{xx}^0 = - \frac{3\tau}{R} \frac{du^0}{dy}, \quad p^0 = -s_{xx}^0 + \frac{3 \operatorname{ctg} \beta}{R} y, \\ u^0 = \frac{1}{6\tau^2} \ln \frac{1 + \sqrt{1 - 36\tau^2}}{1 + \sqrt{1 - 36\tau^2 y^2}} + \frac{1}{6\tau^2} (\sqrt{1 - 36\tau^2 y^2} - \sqrt{1 - 36\tau^2}), \\ \frac{du^0}{dy} = \frac{-1 + \sqrt{1 - 36\tau^2 y^2}}{6\tau^2 y}. \quad (1.4) \end{aligned}$$

which describe the steady flow of a film with a plane stress-free surface, flowing under gravity down an inclined plane, with the no-slip condition at the plane $y = 1$ (Fig. 1).

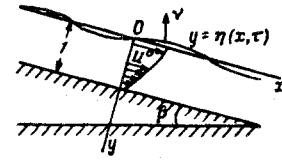


Fig. 1

From (1.4) it can be seen that rectilinear flow of the film is possible when $\tau < 1/6$. When $\tau = 0$, equations (1.4) describe the flow of a film of Newtonian fluid [4].

We expand the velocity $u^0(y, \tau)$ in a power series of the small parameter τ , and truncate the expansion of each expression in (1.4) after the τ^2 term. The result is

$$\begin{aligned} s_{xy}^0 = - \frac{3}{R} y, \quad s_{xx}^0 = -s_{yy}^0 = \frac{9\tau}{R} y^2, \\ u^0 = \frac{3}{2} (1 - y^2) + \frac{27}{4} \tau^2 (1 - y^4), \\ \frac{du^0}{dy} = -3y - 27\tau^2 y^3, \\ p^0 = - \frac{9\tau}{R} y^2 + \frac{3 \operatorname{ctg} \beta}{R} y. \quad (1.5) \end{aligned}$$

2. We proceed to investigate the stability of solution (1.5) with respect to two-dimensional perturbations. Consider a two-dimensional unsteady flow of the form

$$\begin{aligned} u = u^0 + u'(t, x, y), \quad v = v'(t, x, y), \quad p = p^0 + p'(t, x, y), \\ s_{xx} = s_{xx}^0 + s_{xx}'(t, x, y), \quad s_{xy} = s_{xy}^0 + s_{xy}'(t, x, y). \quad (2.1) \end{aligned}$$

Here $u^0, p^0, s_{xx}^0, s_{xy}^0$ are parameters of the basic flow (1.5), and $u', s_{xx}', s_{xy}', v', p'$ are the two-dimensional perturbations.

At the perturbed free surface $y = \eta(t, x)$ we have the kinematic and dynamic conditions [4]

$$\begin{aligned} v = \frac{\partial \eta}{\partial t} + u \frac{\partial \eta}{\partial x}, \\ s_{xx} [\cos^2(v, x) - \cos^2(v, y)] + 2s_{xy} \cos(v, x) \cos(v, y) - p = -S \frac{\partial^2 \eta}{\partial x^2}, \\ 2s_{xx} \cos(v, x) \cos(v, y) + s_{xy} [\cos^2(v, y) - \cos^2(v, x)] = 0, \\ (S = T/\rho d k^2). \quad (2.2) \end{aligned}$$

Here ν is the outer normal to the free surface, and T is the coefficient of surface tension.

At the surface $y = 1$ we have the conditions $u'(t, x, 1) = 0$, $v'(t, x, 1) = 0$. We linearize (1.1), (2.2), taking account of (1.5) and (2.1). Introducing the perturbation stream function by means of the definitions $u' = \partial \Psi / \partial y$, $v' = -\partial \Psi / \partial x$, we represent the perturbations in the form

$$p' = f(y) e^{i\alpha(x-ct)}, \quad s_{xx}' = s_{xx}^*(y) e^{i\alpha(x-ct)},$$

$$s_{xy}' = s_{xy}^*(y) e^{i\alpha(x-ct)}, \quad \Psi = \varphi(y) e^{i\alpha(x-ct)}.$$

The variables $f(y), \varphi(y), s_{xx}^*(y), s_{xy}^*(y)$ are determined by the system

$$\begin{aligned} i\alpha(u^0 - c) \frac{d\varphi}{dy} - i\alpha \frac{du^0}{dy} \varphi = -i\alpha f \left[i\alpha s_{xx}^* + \frac{ds_{xy}^*}{dy} \right], \\ \alpha^2 \varphi (u^0 - c) = - \frac{df}{dy} + i\alpha s_{xy}^* - \frac{ds_{xx}^*}{dy}, \\ R s_{xy}^* = \theta_1 \left(\frac{d^2 \varphi}{dy^2} + \alpha^2 \varphi \right) + 18\tau^2 y^2 \left(\alpha^2 \varphi - \frac{d^2 \varphi}{dy^2} \right) \end{aligned}$$

$$\begin{aligned}
& + \theta_2 \left(6i\alpha y \frac{d\varphi}{dy} - 3i\alpha\varphi \right) + 6\tau^2 \alpha^2 y (u^\circ - c), \\
& R s_{xx}^* = 2\theta_1 i\alpha \frac{d\varphi}{dy} - 6\tau y \frac{d^2\varphi}{dy^2} + \\
& + 9y\tau^2 i\alpha (u^\circ - c) \frac{d^2\varphi}{dy^2} + \frac{27}{2} i\alpha \tau^2 y + 3y i\alpha^3 \tau^2 (u^\circ - c), \\
& \theta_1 = 1 - \tau i\alpha \left[\frac{3}{2} (1 - y^2) - c \right] - \\
& - \tau^2 \left[9y^2 + \frac{9}{4} \alpha^2 (1 - y^2)^2 - 3c\alpha^2 (1 - y^2) + \alpha^2 c^2 \right], \\
& \theta_2 = \tau - \tau^2 i\alpha \left[\frac{3}{2} (1 - y^2) - c \right]. \quad (2.3)
\end{aligned}$$

This system takes into account all terms up to and including τ^2 . The boundary conditions (2.2) yield the relations

$$\begin{aligned}
\eta &= \frac{\varphi(0)}{c - \frac{3}{2}} e^{i\alpha(x-ct)}, \\
2i\alpha \left[\frac{d\varphi}{dy} \right]_0 + 3 \operatorname{ctg} \beta \frac{\varphi(0)}{c - \frac{3}{2}} + f(0)R + \alpha^2 RS \frac{\varphi(0)}{c - \frac{3}{2}} &= 0, \\
\frac{3\varphi(0)}{R(c - \frac{3}{2})} &= s_{xy}^*(0). \quad (2.4)
\end{aligned}$$

Eliminating $f(y)$ from (2.3), we obtain an equation for $\varphi(y)$

$$i\alpha \left[(u^\circ - c) \left(\frac{d^2\varphi}{dy^2} - \alpha^2\varphi \right) - \frac{d^2u^\circ}{dy^2} \varphi \right] = 2i\alpha \frac{ds_{xx}^*}{dy} + \frac{d^2s_{xy}^*}{dy^2} + \alpha^2 s_{xy}^*. \quad (2.5)$$

Here s_{xx}^* and s_{xy}^* have the form given in (2.3).

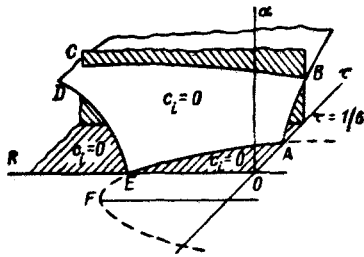


Fig. 2

Using the first equation in (2.3), we eliminate $f(0)$ from the boundary condition (2.4). The boundary condition for equation (2.5) at $y = 0$ is then

$$\begin{aligned}
& \frac{\alpha}{c - \frac{3}{2}} (3 \operatorname{ctg} \beta + \alpha^2 RS) \varphi(0) + \alpha \left\{ 2i\alpha \left[\frac{d\varphi}{dy} \right]_0 + \right. \\
& \left. + R \left(c - \frac{3}{2} \right) \left[\frac{d\varphi}{dy} \right]_0 + R s_{xy}^*(0) \right\} - \\
& - iR \left[\frac{ds_{xy}^*}{dy} \right]_0 = 0, \quad \frac{3\varphi(0)}{R(c - \frac{3}{2})} = s_{xy}^*(0). \quad (2.6)
\end{aligned}$$

The conditions $u'(t, x, 1) = v'(t, x, 1) = 0$ yield

$$\varphi(1) = \left[\frac{d\varphi}{dy} \right]_1 = 0. \quad (2.7)$$

The problem of solving (2.5) with the boundary conditions (2.6), (2.7) leads, as is well known [4], to the problem of finding $c = c(R, \alpha, \beta, \tau)$.

3. To find $c = c(R, \alpha, \beta, \tau)$ for small α we expand φ and c in terms of the parameter α :

$$\varphi = \varphi^\circ + \alpha\varphi' + \alpha^2\varphi'' + \dots, \quad c = c^\circ + \alpha c' + \alpha^2 c'' + \dots \quad (3.1)$$

Substitute (3.1) into (2.5)–(2.7) and (2.3). For $\alpha = 0$ we obtain

$$\begin{aligned}
& \frac{d^2}{dy^2} \left[(1 - 27\tau^2 y^2) \frac{d^2\varphi^\circ}{dy^2} \right] = 0, \quad \left[\frac{d^2\varphi^\circ}{dy^2} \right]_0 = 0, \\
& \frac{3\varphi^\circ(0)}{c^\circ - \frac{3}{2}} = \left[\frac{d^2\varphi^\circ}{dy^2} \right]_0, \quad \varphi^\circ(1) = \left[\frac{d\varphi^\circ}{dy} \right]_1 = 0. \quad (3.2)
\end{aligned}$$

From (3.2) we find [4] the solution to the zeroth approximation

$$c^\circ = 3 + \frac{81}{4}\tau^2, \quad \varphi^\circ = (1 - y)^2 + \tau^2 (9y - 27y^2 + \frac{9}{4}y^4). \quad (3.3)$$

The variable $\varphi'(y)$ is then determined by the equation

$$\frac{d^2}{dy^2} \left[(1 - 27\tau^2 y^2) \frac{d^2\varphi'}{dy^2} \right] = iR \left[(u^\circ - c^\circ) \frac{d^2\varphi^\circ}{dy^2} - \frac{d^2u^\circ}{dy^2} \varphi^\circ \right], \quad (3.4)$$

with the boundary conditions

$$\begin{aligned}
& \frac{1}{c^\circ - \frac{3}{2}} [3 \operatorname{ctg} \beta + \alpha^2 RS] \varphi^\circ(0) + \\
& + R \left(c^\circ - \frac{3}{2} \right) \left[\frac{d\varphi^\circ}{dy} \right]_0 - i \left[\frac{d^2\varphi^\circ}{dy^2} \right]_0 + 3\tau \left[\frac{d\varphi^\circ}{dy} \right]_0 = 0, \\
& - \frac{3\varphi^\circ(0)}{c^\circ - \frac{3}{2}} + \frac{3\varphi^\circ(0)c'}{(c^\circ - \frac{3}{2})^2} + \left[\frac{d^2\varphi'}{dy^2} \right]_0 = 0, \quad \varphi'(1) = \left[\frac{d\varphi'}{dy} \right]_1 = 0. \quad (3.5)
\end{aligned}$$

Integrating (3.4), taking into account (3.3) and the boundary conditions (3.5), we find

$$\begin{aligned}
c' &= i(1 + \frac{27}{4}\tau^2) \left[\frac{6}{5}R(1 + 9.12\tau^2) - \right. \\
& \left. - (1 - 3.30\tau^2) \operatorname{ctg} \beta + 3\tau - \frac{1}{3}RS(1 - 3.30\tau^2) \right]. \quad (3.6)
\end{aligned}$$

Here, as everywhere else, we neglect terms of orders higher than τ^2 .

For small α we retain the two leading terms in the expansions. We obtain then $c = c^\circ + \alpha c'$, where c° and c' are given by (3.3) and (3.6). Equating the imaginary part of c with zero, we find

$$\begin{aligned}
& \alpha \left[\frac{6}{5}R(1 + 9.12\tau^2) - (1 - 3.30\tau^2) \operatorname{ctg} \beta + \right. \\
& \left. 3\tau - \frac{1}{3}\alpha^2 RS(1 - 3.30\tau^2) \right] = 0. \quad (3.7)
\end{aligned}$$

For $\beta = \text{const}$, Eq. (3.7) defines a neutral surface in the R, α, M space. This surface consists of the plane $\alpha = 0$ and the surface ABCDE, which intersect along the parabola (Fig. 2)

$$R = \frac{5}{6} \operatorname{ctg} \beta (1 - 10.35\tau^2) - 2.50\tau. \quad (3.8)$$

The portion EA of the parabola (3.8) is a branching line of the neutral surface.

For $\tau = \lambda = 0$ we have in the α, R plane a neutral curve, consisting of the axis $\alpha = 0$ and the line ED. The branch point E defines then [4] the minimal critical value $R = 5/6 \operatorname{ctg} \beta$ for $S \neq 0$.

From (3.7) one can see that when $S \neq 0$ the inequality

$$R \leq \frac{5}{6} (1 - 10.35\tau^2) \operatorname{ctg} \beta - 2.50\tau$$

must hold on the surface ABCDE if α is to be real. Therefore for small α the surface ABCDE cannot intersect the α axis.

Taking account of (3.6) one can easily see that when $\tau < 1/6$, $S \neq 0$, the values of c_i are negative for $\alpha > \alpha_0$ (α_0 are the values of α on the surface ABCDE) and positive for $\alpha < \alpha_0$.

Taking into account that $\lambda = \mu/G$, where G is the elastic shear modulus, we find from (1.3) that τ is independent of μ . Therefore, if $\tau = 0$ because of $G = \infty$, then for $R = 0$, $\mu = \infty$ we find [4] that the surface ABCDE does not intersect the α axis in the case $S \neq 0$. For $S = 0$ the surface ABCDE intersects the α axis at $\alpha = \infty$.

For a Newtonian fluid ($G = \infty$) the branch point E of the neutral curve coincides with the origin for $\beta = \pi/2$. For a viscoelastic fluid with $G = \text{const} \neq \infty$, the Reynolds number is equal to zero at the branch point of the neutral curve when

$$\tau = A = 1/17.25 [-2.50 + (6.25 + 2.88 \operatorname{ctg}^2 \beta)^{1/2}] \operatorname{tg} \beta.$$

Consequently, in the case of a Newtonian fluid the critical Reynolds number is equal to zero for flow on a vertical plane ($\beta = \pi/2$), whereas in the case of a viscoelastic fluid the critical Reynolds number is equal to zero for $\beta = \arccos(\operatorname{ctg} \beta)$. Thus, for example, for $\tau = 0.1$ the critical Reynolds number is equal to zero for flow along a plane inclined at $\beta \approx 73^\circ$ with respect to the horizontal.

Fig. 2 shows the neutral surface for the case $S \neq 0$. The equations of motion (1.3) have an exact solution, corresponding to rectilinear laminar flow, for $\tau < 1/6$, and this flow is stable with respect

to two-dimensional perturbations in the region between the planes $\alpha = 0$, $\tau = 0$, $R = 0$ and the surface ABCDE. The plane $\alpha = 0$ and the surface ABCDE are a neutral surface, whose branching line EA is an arc of the parabola (3.8) in the plane $\alpha = 0$. The parabola has an axis parallel to the R axis, and its vertex is at the point F, whose coordinates are

$$\alpha = 0, \quad R = 0.83 \operatorname{ctg} \beta + 0.18 \operatorname{tg} \beta, \quad \tau = -0.15 \operatorname{tg} \beta.$$

The flow of a film of viscoelastic fluid is less stable than the flow of a Newtonian fluid with the same viscosity.

REFERENCES

1. L. I. Sedov, "Different definitions of the rate of change of a tensor," PMM, vol. 24, no. 3, 1960.
2. J. G. Oldroyd, "Non-Newtonian effects in steady motion of some idealized elastico-viscous liquids," Proc. Roy. Soc. A, vol. 245, no. 1241, p. 278-297, 1958.
3. Yu. P. Ivanilov, "On the stability of plane-parallel flow of a viscous fluid over an inclined bottom," PMM, vol. 24, no. 2, 1960.
4. Yih Chia-shun, "Stability of liquid flow down an inclined plane," Phys. Fluids, vol. 6, no. 3, 1963.

11 December 1964

Voronezh